

# Vacuum-polarization corrections to the PNC amplitude in $^{133}\text{Cs}^*$

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## Abstract

One-loop radiative corrections to atomic wave functions associated with vacuum polarization in the nuclear Coulomb field are evaluated for the  $6s - 7s$  PNC amplitude in  $^{133}\text{Cs}$ . These corrections increase the size of the PNC amplitude by 0.4%.

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\*Phys. Rev. Lett. 87, 233001 (2001)

## Background - $^{133}\text{Cs}$

1.  $E_{\text{PNC}} = F \times \frac{Q_W}{-N}$  evaluated (1%) by two groups.<sup>1</sup>
2.  $E_{\text{PNC}}/\beta$  precisely measured.<sup>2</sup>
3.  $\beta$  measured and theory error reviewed (0.4%):<sup>3</sup>  
 $\Delta Q_W = Q_W^{\text{expt}} - Q_W^{\text{SM}} = 2.5 \sigma$
4. Breit interaction reconsidered:<sup>4</sup>  $2.5 \sigma \rightarrow 1 \sigma$
5. Fact that bound-state radiative corrections are comparable to Breit corrections pointed out.<sup>5</sup>

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<sup>1</sup>V. A. Dzuba *et al.*, Phys. Lett. A **141**, 147 (1989); S. A. Blundell *et al.*, Phys. Rev. D **45**, 1602 (1992).

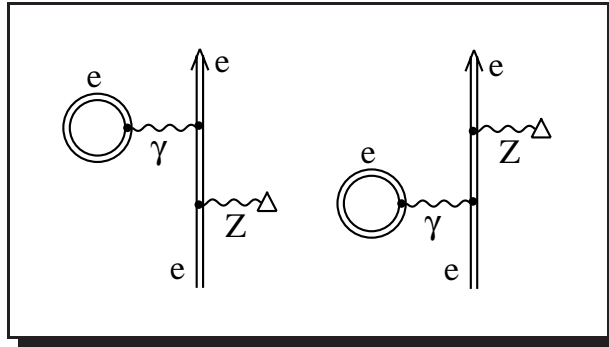
<sup>2</sup>C. S. Wood *et al.*, Science **275**, 1759 (1997).

<sup>3</sup>S. C. Bennett and C. E. Wieman, Phys. Rev. Lett. **82**, 2484 (1999).

<sup>4</sup>A. Derevianko, Phys. Rev. Lett. **85**, 1618 (2000).

<sup>5</sup>O. P. Sushkov, Phys. Rev. A **63**, 042504 (2001).

## Which Corrections?



The double line represents the electron in the field of the nucleus. To leading order in powers of  $Z\alpha$ ,<sup>6</sup>

$$-\frac{Z}{r} \Rightarrow -\frac{Z}{r} + \delta V$$

where  $\delta V$  given by the Uehling potential<sup>7</sup>:

$$\delta V(r) = -\frac{2\alpha Z}{3\pi r} \int_1^\infty dt \sqrt{t^2 - 1} \left( \frac{1}{t^2} + \frac{1}{2t^4} \right) e^{-2ctr}$$

<sup>6</sup>E. H. Wichmann and N. H. Kroll, Phys. Rev. **101**, 843 (1956)

<sup>7</sup>E. A. Uehling, Phys. Rev. **48**, 55 (1935)

## Distributed Nuclear Charge

The nuclear charge distribution for  $^{133}\text{Cs}$  is

$$\rho(r) = \frac{\rho_0}{1 + \exp [(r - c_{\text{nuc}})/a_{\text{nuc}}]}$$

with  $c_{\text{nuc}} = 5.675$  fm and 10%–90% fall-off distance is  $t_{\text{nuc}} = 2.3$  fm, corresponding to  $a_{\text{nuc}} = 0.523$  fm and to a root-mean-square radius of the nuclear charge distribution  $R_{\text{rms}} = 4.807$  fm.

For a distributed charge:<sup>8</sup>

$$\delta V(r) = -\frac{2\alpha^2}{3r} \int_0^\infty dx x \rho(x) \int_1^\infty dt \sqrt{t^2 - 1} \times \left( \frac{1}{t^3} + \frac{1}{2t^5} \right) \left( e^{-2ct|r-x|} - e^{-2ct(r+x)} \right).$$

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<sup>8</sup>L. W. Fullerton and G. A. Rinker, Jr., Phys. Rev. A **13**, 1283 (1976)

## Test of Method

Comparison of corrections to inner-shell Dirac-Coulomb energies  $\delta E$  on replacing  $V \rightarrow V + \delta V$  with results from QED perturbation theory<sup>9</sup>.

State	Present	Perturbation Theory			Total
	$\delta E$	Uehl	Uehl-FS	FS	
$1s_{1/2}$	-0.1419	-0.2584	0.0010(1)	0.1159	-0.1415(1)
$2s_{1/2}$	-0.1554	-0.2901	0.0011(1)	0.1339(1)	-0.1551(2)
$2p_{1/2}$	-0.0100	-0.0145	0.0000(1)	0.0045	-0.0100(1)
$2p_{3/2}$	-0.0016	-0.0016	0.0000	0.0000	-0.0016

Units:  $\alpha mc^2(\alpha Z)^4/(\pi n^3)$

<sup>9</sup>W. R. Johnson and G. Soff, At. Data Nucl. Data Tables **33**, 405 (1985)

# Calculation #1 - Weak Hartree Fock

## No Correlation Corrections

Effect of PNC:  $\phi_v^{\text{HF}} \rightarrow \phi_v^{\text{HF}} + \delta\phi_v^{\text{HF}}$

$$(h_0 + V^{\text{HF}} - \epsilon_v^{\text{HF}}) \delta\phi_v^{\text{HF}} = -h_{\text{PNC}} \phi_v^{\text{HF}}$$

$$E_{\text{PNC}} = \langle \phi_{7s}^{\text{HF}} | D | \delta\phi_{6s}^{\text{HF}} \rangle + \langle \delta\phi_{7s}^{\text{HF}} | D | \phi_{6s}^{\text{HF}} \rangle$$

Corrections to  $E_{\text{PNC}}$  amplitude for  $^{133}\text{Cs}$

Units:  $iea_0 \times 10^{-11} (-Q_W/N)$

Type	$\langle \phi_{7s}   D   \delta\phi_{6s} \rangle$	$\langle \delta\phi_{7s}   D   \phi_{6s} \rangle$	$E_{\text{PNC}}$
HF	-2.7492	1.0144	-0.7395
HF + $\delta V$	-2.7604	1.0186	-0.7425
$\Delta$ (%)	0.41	0.41	0.41

(This calculation gives a value of the PNC amplitude that is 20% smaller than the fully correlated value.)

## Calculation #2 - Weak RPA

### With Correlation Corrections

$$(h_0 + V^{\text{HF}} - \epsilon_v^{\text{HF}}) \delta\phi_v^{\text{RPA}} = - [h_{\text{PNC}} + V_{\text{PNC}}^{\text{HF}}] \phi_v^{\text{HF}},$$

$$E_{\text{PNC}} = \langle \phi_{7s}^{\text{HF}} | D | \delta\phi_{6s}^{\text{RPA}} \rangle + \langle \delta\phi_{7s}^{\text{RPA}} | D | \phi_{6s}^{\text{HF}} \rangle$$

Corrections to  $E_{\text{PNC}}$  amplitude for  $^{133}\text{Cs}$ , with RPA.

Units:  $iea_0 \times 10^{-11} (-Q_W/N)$

Type	$\langle \phi_{7s}   D   \delta\phi_{6s} \rangle$	$\langle \delta\phi_{7s}   D   \phi_{6s} \rangle$	$E_{\text{PNC}}$
RPA	-3.4570	1.2726	-0.9269
RPA+ $\delta V$	-3.4712	1.2778	-0.9307
$\Delta$ (%)	0.41	0.41	0.41

(This calculation gives a value of the PNC amplitude that is 3% larger than the fully correlated value.)

Conclusion:

VP corrections increase the PNC amplitude by 0.41%.

## Comments on the PV Hamiltonian<sup>10</sup>

$$H_{\text{PV}} = \frac{G_{\mu}}{\sqrt{2}} \left[ \bar{e} \gamma_{\mu} \gamma_5 e \left( C_{1u} \bar{u} \gamma_{\mu} u + C_{1d} \bar{d} \gamma_{\mu} d + \dots \right) \right. \\ \left. + \bar{e} \gamma_{\mu} e \left( C_{2u} \bar{u} \gamma_{\mu} \gamma_5 u + C_{2d} \bar{d} \gamma_{\mu} \gamma_5 d + \dots \right) \right]$$

where  $\dots$  corresponds to  $q = s, t, b, c$

$$C_{1u} = \left( -\frac{1}{2} + \frac{4}{3} s^2 \right)$$

$$C_{1d} = \left( \frac{1}{2} - \frac{2}{3} s^2 \right)$$

$$C_{2u} = -\frac{1}{2} (1 - 4s^2)$$

$$C_{2d} = \frac{1}{2} (1 - 4s^2)$$

$$s^2 = \sin^2 \theta_W$$

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<sup>10</sup>W. J. Marciano in *Precision Tests of the Standard Electroweak Model*, Ed. P. Langacker, (World Scientific, Singapore, 1995), p. 170



## Electroweak Radiative Corrections

For atomic PNC with nonrelativistic nucleons, the Hamiltonian (in the electron sector) reduces to

$$h_{\text{PNC}} = \frac{G_{\mu}}{\sqrt{2}} Q_W \gamma_5 \rho_{\text{nuc}}(r)$$

with

$$Q_W = 2[(2Z + N)C_{1u} + (Z + 2N)C_{1d}]$$

Using  $s^2 = 0.23114(17)$ , one finds

$$Q_W^{\text{SM}} = -73.86 \pm 0.03 \text{ (no radiative corrections)}$$

## Electroweak Radiative Corrections

With electroweak radiative corrections,<sup>11</sup> the coupling constants can be written

$$C_{1u} = \rho'_{\text{PV}} \left( -\frac{1}{2} + \frac{4}{3} \kappa'_{\text{PV}} s^2 \right) + \lambda_{1u}$$

$$C_{1d} = \rho'_{\text{PV}} \left( \frac{1}{2} - \frac{2}{3} \kappa'_{\text{PV}} s^2 \right) + \lambda_{1d}$$

For Atomic PNC:

$$\rho'_{\text{PV}} = 0.9878$$

$$\kappa'_{\text{PV}} = 1.0026$$

$$\lambda_{1u} = -1.9 \times 10^{-5}$$

$$\lambda_{1d} = 3.7 \times 10^{-5}$$

$$Q_W^{\text{SM}} = -73.09 \pm 0.03 \text{ (with radiative corrections)}$$

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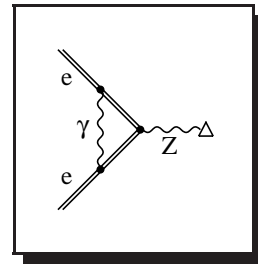
<sup>11</sup>D. E. Groom et al., Euro. Phys. J. C **15**, 1 (2000)

## $\alpha Z$ -dependent terms

Perturbation Theory:<sup>12</sup>

$$\rho'_{\text{PV}} = 1 - \frac{\alpha}{2\pi} + \frac{\alpha(m_Z)}{2\pi} \left\{ \frac{3}{8s^2} \frac{m_t^2}{m_W^2} + \frac{3}{8s^4} \ln c^2 - \frac{7}{8s^2} + \frac{3\xi}{8s^2} \left( \frac{\ln(c^2/\xi)}{c^2 - \xi} + \frac{1}{c^2} \frac{\ln \xi}{1 - \xi} \right) - \frac{1}{s^2} + \dots \right\}$$

The term  $-\frac{\alpha}{2\pi}$  is an electromagnetic vertex correction that should be evaluated using bound-state electron propagators.



Assume that the binding corrections modify

$$-\frac{\alpha}{2\pi} \rightarrow -\frac{\alpha}{2\pi} \pm \frac{\alpha}{2\pi},$$

and we find  $Q_W^{\text{SM}} = -73.09 \rightarrow -73.09 \pm 0.08$

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<sup>12</sup>W. J. Marciano in *Precision Tests of the Standard Electroweak Model*, Ed. P. Langacker, (World Scientific, Singapore, 1995), p. 182

## Conclusion

Using the average of the two most accurate theoretical values (plus Breit) of  $E_{\text{PNC}} \times 1.0041$ :

$$E_{\text{PNC}} = -9.057 \pm 0.037 \text{ } iea_0 \times 10^{-12} (-Q_W/N),$$

where we use the 0.4% error estimate of Bennett and Wieman.<sup>13</sup> Combining this with the experimental value of  $E_{\text{PNC}}/\beta^{14}$  leads to an experimental value for the weak charge

$$Q_W^{\text{expt}}(^{133}\text{Cs}) = -72.12 \pm (0.28)_{\text{expt}} \pm (0.34)_{\text{theor}}$$

This differs by **2.2  $\sigma$**  from the standard-model value<sup>15</sup>,

$$Q_W^{\text{SM}}(^{133}\text{Cs}) = -73.09 \pm 0.03$$

<sup>13</sup>S. C. Bennett and C. E. Wieman, Phys. Rev. Lett. **82**, 2484 (1999); **82**, 4153 (1999); **83**, 889 (1999)

<sup>14</sup>C. S. Wood *et al.*, Science **275**, 1759 (1997)

<sup>15</sup>D. E. Groom *et al.*, Euro. Phys. J. C **15**, 1 (2000).