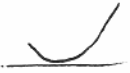


Physics 131 Spring 2001

Final Exam Solutions

Multiple Choice

1. (B) The net force is just that of the four springs combined, so $F = 4kx = k'x$, $\omega = \sqrt{k'/m} = 2\sqrt{k/m}$
2. (C) For circular orbits, $\frac{mv^2}{r} = \frac{GMm}{r^2}$, $v = \sqrt{\frac{GM}{r}}$, so the smallest orbit has the largest tangential speed.
3. (C) The torques must be equal for static equilibrium:
 $\sum \tau = 0 = M_1 R_1 g - M_2 R_2 g \rightarrow M_1 R_1 = M_2 R_2$
4. (E) They will bounce back in the opposite directions, but with smaller speeds: momentum is conserved ($\sum p_x = 0$), but energy is not.
5. (E) The total work around the loop must be zero, and, since the work should only depend on the endpoints, $W_1 = -W_2$.
6. (C) The easiest way to see this is to draw them:
 A. $U = bx^2$  SHM ✓ B. $U = \sin(bx - \pi/2)$  ✓ SHM
 C. $U = \cos(bx)$  (X) SHM D. $U(x) = e^x - e^{-x}$  ✓ SHM
 E. $U(x) = x^2 + bx^3$  ✓ SHM
7. (D) Work done against the drag force = $-F_{\text{drag}} \cdot \text{distance}$.
 $F_1 \propto (50 \text{ mph})^2 = b(50 \text{ mph})^2$, $F_2 = b(70 \text{ mph})^2$
 $d_1 = (50 \text{ mph})(10 \text{ h}) = 500 \text{ miles}$, $d_2 = 70 \cdot 10 = 700 \text{ miles}$.
 $\frac{W_2}{W_1} = \frac{F_2 \cdot d_2}{F_1 \cdot d_1} = \frac{b(70)^2(700)}{b(50)^2(500)} = 2.74$, so $2.74 \times 20 \text{ gal} \approx 55 \text{ gal}$.

Multiple Choice, cont.

8. © The wave on the right has the same amplitude, twice the frequency (or half the wavelength), and is shifted in phase. To halve the wavelength, $k' = \frac{2\pi}{\lambda'} = 2 \left(\frac{2\pi}{\lambda} \right) = 2k$. Choice C is the only one possible.

Problems

II. A.) Conserve Energy: $E_i = E_f : \frac{1}{2}mv^2 = \frac{1}{2}kx^2$, $x = v\sqrt{\frac{m}{k}} = 0.45\text{ m}$

B.) $\omega = \sqrt{k/m}$, $f = \frac{\omega}{2\pi} = \frac{\sqrt{k/m}}{2\pi} = 0.711\text{ sec}^{-1} = \underline{0.711\text{ Hz}}$

C.) $A = x_{\max} = \underline{0.45\text{ m}}$

D.) This is $1/4$ of an oscillation, so it takes $T/4\text{ sec} = \frac{1}{4f} = \underline{0.35\text{ sec}}$.
(If you said $3/4$ of an oscillation, $3T/4 = 1.05\text{ sec}$)

- III. A.) Since there is no external force in the horizontal plane, P_x is conserved. Since gravity acts in the vertical plane, P_y is not conserved, even if $P_y^i = P_y^f = 0$.

- B.) Conserve energy and momentum:

$$E_i = E_f$$

$$mgh = \frac{1}{2}mU^2 + \frac{1}{2}MV^2$$

$$P_x^i = P_x^f$$

$$0 = mU - MV$$

$$\Rightarrow mU = MV, \quad V = \frac{m}{M}U$$

assume V is in $-\hat{i}$ direction

$$\Rightarrow mgh = \frac{1}{2}mU^2 + \frac{1}{2}M\left(\frac{m}{M}U\right)^2, \quad 2mgh = mU^2 + \frac{m^2}{M}U^2,$$

$$U^2\left(1 + \frac{m}{M}\right) = 2gh, \quad U^2 = \frac{2ghM}{m+M}, \quad \boxed{U = \sqrt{\frac{2ghM}{m+M}} \hat{i}}$$

$$V = \frac{m}{M}U = \sqrt{\frac{2ghm^2M}{M^2(M+m)}} = \sqrt{\frac{2ghm^2}{M^2+Mm}} \Rightarrow \boxed{V = -\sqrt{\frac{2ghm^2}{M^2+Mm}} \hat{i}}$$

- C.) Since friction is an internal force, P_x is still conserved. The final velocities U and V will just be smaller.

Problems, (cont.)

IV. A.) $\bar{v} = \frac{\Delta x}{\Delta t} = \frac{2011 \text{ m}}{119.8 \text{ s}} = \boxed{16.8 \text{ m/s}}$

B.) If $a = \text{const.}$, $v = at$, $\bar{v} = \frac{at_f - 0}{2} = a \frac{t_f}{2} = \frac{1}{2} v_f$

So, $\bar{v} = \frac{1}{2} v_f$, $\boxed{v_f = 33.6 \text{ m/s}}$

C.) $P = \Delta KE / \Delta t$, so $P \Delta t = \Delta KE = \frac{1}{2} m v_f^2 - 0$.

$\frac{1}{2} m v_f^2 = (746 \text{ W})(119.8 \text{ s}) = 89,300 \text{ J}, \Rightarrow v^2 = 510 \text{ m}^2/\text{s}^2$

$\boxed{v = 22.6 \text{ m/s}}$

D.) At constant power, $P = F \cdot v = m a v$, so the acceleration is $a = \frac{P}{m v}$. Since v is increasing, the acceleration will decrease with time. Constant acceleration requires a constant force, which you can see from $P = F v$ is impossible for constant power output. Another way to see this is that power $= \Delta KE / \Delta t$. Since $\Delta KE \propto v^2$, it takes smaller and smaller increases in v (as v gets larger) to produce the same ΔKE , so a decreases as v gets larger.

V. A.) $I_{\text{sun}} = \frac{2}{5} M_{\text{sun}} R^2$, $L_{\text{sun}} = I_{\text{sun}} \omega_{\text{sun}}$, $\omega_{\text{sun}} = \frac{2\pi}{T_{\text{sun}}} = 2.87 \times 10^{-6} \text{ (1/s)}$
 $= 3.92 \times 10^{47} \text{ kg m}^2$ $\rightarrow \underline{L_{\text{sun}} = 1.13 \times 10^{42} \text{ kg m}^2/\text{s}}$

B.) $I_{\text{RG}} = \frac{2}{5} M_{\text{RG}} R_{\text{RG}}^2 = 8 \times 10^{51} \text{ kg m}^2$,

Angular momentum will be conserved, so $L_{\text{sun}} = L_{\text{RG}} = I_{\text{RG}} \omega_{\text{RG}}$

$\omega_{\text{RG}} = 1.41 \times 10^{-10} \text{ rad/s} = \frac{2\pi}{T_{\text{RG}}}$, $\underline{T_{\text{RG}} = 4.4 \times 10^{10} \text{ s} = 5.15 \times 10^5 \text{ days}}$

C.) The outer shell blows away, leaving a smaller star rotating at the same angular velocity

$I_{\text{new}} = \frac{2}{5} M_{\text{new}} R_{\text{new}}^2 = \frac{2}{5} (8.4 \times 10^{29} \text{ kg}) (0.75 \times 10^{10})^2 = 1.89 \times 10^{51} \text{ kg m}^2$
 (cont.)

Problems (cont.)

V. C) (cont.) Incidentally, $\frac{M_{\text{new}}}{M_{\text{tot}}} = \frac{V_{\text{new}}}{V_{\text{tot}}} = \frac{\frac{4}{3}\pi R_{\text{new}}^3}{\frac{4}{3}\pi R_{\text{RG}}^3} = (0.75)^3 = 0.42$

So, $L_{\text{new}} = I_{\text{new}} \omega_{\text{RG}} = (1.89 \times 10^{51})(1.41 \times 10^{-10}) = \underline{2.67 \times 10^{41} \text{ kg m}^2/\text{s}}$

D.) $I_{\text{WD}} = (2/5)(M_{\text{new}})(R_{\text{WD}})^2 = 1.21 \times 10^{43} \text{ kg m}^2.$

Since angular momentum will be conserved in the collapse,
 $L_{\text{WD}} = L_{\text{new}} = I_{\text{WD}} \omega_{\text{WD}}, \omega_{\text{WD}} = 0.022 \text{ rad/s}$

$T_{\text{WD}} = \frac{2\pi}{\omega_{\text{WD}}} = \underline{284 \text{ seconds}} = 4.7 \text{ minutes!}$

E.) At the equator, $v = \frac{2\pi R_{\text{WD}}}{T_{\text{WD}}} = \underline{1.33 \times 10^5 \text{ m/s}} \approx 300,000 \text{ mph!}$

VI. A.) $\rho = \frac{M}{V} = \frac{8.4 \times 10^{29} \text{ kg}}{\frac{4}{3}\pi R_{\text{WD}}^3} = 9.3 \times 10^8 \text{ kg/m}^3 \quad (84,400 \times \rho_{\text{Lead}})$

one $\text{cm}^3 = (0.01)^3 \text{ m}^3$, so $m = \rho V = (9.3 \times 10^8)(1 \times 10^{-6})$
 $= \underline{930 \text{ kg.}}$

B.) Weight = $m\vec{g} = m \left(\frac{GM_{\text{WD}}}{R_{\text{WD}}^2} \right) = 930 \text{ kg} \left(\frac{(6.67 \times 10^{-11})(8.4 \times 10^{29})}{(6 \times 10^6)^2} \right)$
 $= 1.45 \times 10^9 \text{ N} \quad (325 \text{ million lbs. !})$

C.) For escape velocity, $E_{\text{tot}} = 0 = \frac{1}{2}mv_{\text{esc}}^2 - \frac{GM_{\text{WD}}m}{R_{\text{WD}}}$
 $v_{\text{esc}} = \sqrt{\frac{2GM_{\text{WD}}}{R_{\text{WD}}}} = 4.32 \times 10^6 \text{ m/s}, \text{ about } 1\% \text{ of the speed of light.}$